# **Rapid Communication**

# Chemical signatures of Buckminsterfullerene, $C_{60}$ , under chemical ionization conditions

Detlef Schröder, Diethard K. Bohme, Thomas Weiske and Helmut Schwarz Institut für Organische Chemie der Technischen Universität Berlin, Strasse des 17 Juni 135, W-1000 Berlin 12 (Germany)

(Received 22 May 1992)

#### ABSTRACT

Chemical ionization (CI) spectra are reported for the Buckminsterfullerene molecule,  $C_{60}$ , with the CI reagents hydrogen, ammonia, water, hydrogen chloride, methane, ethylene and isobutane. The spectra are characterized by a strong peak for protonated  $C_{60}$ ,  $C_{60}$  H<sup>+</sup>, except with ammonia. The spectra with hydrogen and ammonia are striking in that they show extensive hydrogenation of  $C_{60}^{+}$ .  $C_{60}$  H<sub>x</sub><sup>+</sup> peaks up to  $C_{60}$  H<sub>37</sub><sup>+</sup> are observed in hydrogen, while the ammonia spectrum shows a preference for an odd number of hydrogen atoms below x = 13. The spectra with methane, ethylene and isobutane show complicated patterns of peaks corresponding to the derivatization of  $C_{60}^{++}$  with  $nCH_x$  units, where *n* has values up to n = 32 in the case of methane. Ion/molecule reactions are discussed which may lead to the formation of  $C_{60}^{++}$ , the further hydrogenation of this ion, and the derivatization of  $C_{60}^{++}$  more generally. The results in general indicate a remarkable ability for  $C_{60}$  to become hydrogenated and derivatized under CI conditions.

Keywords: fullerenes; chemical ionization; ions; ion/molecule reactions.

# INTRODUCTION

The chemical reactivity of Buckminsterfullerene,  $C_{60}$ , both in solution and in the gas phase, has become of great interest since the availability of this compound in bulk quantities [1]. Since the first detection of its cation in laser vaporization experiments with graphite, neutral  $C_{60}$  has been proposed as an important interstellar molecule [2]. The failure to detect  $C_{60}$  in interstellar environments [3] has shifted interest to the  $C_{60}$  cation [4] and its derivatives, particularly  $C_{60}H^+$  [5]. The few experimental studies which have directly explored the chemical reactivity of  $C_{60}$  (for a review see ref. 6) show that this species is quite unreactive near thermal energies. For example,  $C_{60}^{*+}$  has been

Correspondence to: H. Schwarz, Institut für Organische Chemie der Techischen Universität Berlin, Strasse des 17 Juni 135, W-1000 Berlin 12, Germany.

<sup>&</sup>lt;sup>1</sup>Permanent address: Department of Chemistry and Centre for Research in Earth and Space Science, York University, North York, Ont. M3J 1P3, Canada.

shown to be unreactive with H<sub>2</sub>, O<sub>2</sub>, NO, NH<sub>3</sub> and C<sub>4</sub>H<sub>2</sub> at low pressures [7], while selected-ion flow tube (SIFT) studies at 0.40 Torr in helium have shown that C<sub>60</sub><sup>++</sup> and C<sub>70</sub><sup>++</sup> are unreactive,  $k < 1 \times 10^{-12}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>, with a wide variety of other molecules including CO, CH<sub>4</sub>, H<sub>2</sub>O, C<sub>2</sub>H<sub>6</sub>, C<sub>2</sub>H<sub>2</sub>, C<sub>2</sub>H<sub>4</sub> and *i*-C<sub>4</sub>H<sub>8</sub> [8]. However, addition reactions have been observed with ammonia and several aliphatic amines [9]. To date, the only bimolecular reactions of C<sub>60</sub><sup>++</sup> which have been reported are charge transfer reactions with neutral molecules, X, having IE(X) < IE(C<sub>60</sub>) = 7.61 ± 0.11 eV [10]. Since C<sub>60</sub><sup>++</sup> displays so little reactivity with molecules, its reactivity with atoms and radicals becomes interesting, especially in view of the radical-sponge behaviour seen for neutral C<sub>60</sub> with the species H<sup>+</sup> [11], F<sup>+</sup> [12], O [13,14], OH<sup>+</sup> [14], CH<sub>2</sub> [14], CH<sub>3</sub> [14–16], CF<sub>3</sub> [16], (CH<sub>3</sub>)<sub>3</sub>CO<sup>-</sup> [16], C<sub>6</sub>H<sub>5</sub> [15], C<sub>6</sub>H<sub>5</sub>CH<sub>2</sub><sup>+</sup> [16], C<sub>6</sub>H<sub>5</sub>S<sup>+</sup> [16] and C(C<sub>6</sub>H<sub>5</sub>)<sub>2</sub>[17]. Also, a very recent SIFT study has shown that C<sub>60</sub><sup>++</sup> will attach H atoms in sequential addition reactions [18].

Less is known of reactions of molecular ions with neutral  $C_{60}$ , although proton transfer and ion addition reactions have been reported in studies of the proton affinity of  $C_{60}$  [19]. Atomic ions have been shown either to add to  $C_{60}$ or transfer a charge, depending on the recombination energy of the atomic ion [20]. Here we report results of experiments in which the vapour of  $C_{60}$  is introduced into a conventional chemical ionization (CI) source along with a CI gas, viz. H<sub>2</sub>, CH<sub>4</sub>, NH<sub>3</sub>, H<sub>2</sub>O, HCl, C<sub>2</sub>H<sub>4</sub> and *i*-C<sub>4</sub>H<sub>10</sub>. The resulting spectra show a remarkable ability for C<sub>60</sub> cations to become derivatized and point a substantial chemistry for fullerene molecules under CI conditions.

#### **EXPERIMENTAL**

The experiments were performed with a modified VG-ZAB tandem mass spectrometer of BEBE configuration (B stands for magnetic and E for electric sector), which has been described in detail elsewhere [21]. CI of C<sub>60</sub> was afforded by use of different reagent gases in a CI source (electron energy 100 eV; repeller voltage  $\approx 0$  V; source housing pressure  $\approx 10^{-4}$  mbar). Source spectra were taken at a mass resolution  $m/\Delta m$  of 2000 and the mass scale was calibrated with (Cs<sub>x</sub>I<sub>x-1</sub>)<sup>+</sup> cluster ions. C<sub>60</sub> was introduced into the source as a C<sub>60</sub>: C<sub>70</sub> mixture ( $\approx 90:10$ ) via the solid probe inlet and heated to 480– 550°C. The temperature of the ion source was increased to 300°C to reduce condensation of fullerenes in the source block<sup>a</sup>. The spectra were accumulated and processed on line with the VG 11/250 data system.

<sup>&</sup>lt;sup>a</sup> However, the large excess of reagent gas led to a serious coating of the source block with fullerenes so that frequent cleaning and maintenance of the ion source was necessary.



Fig. 1. CI spectra for  $C_{60}$ . From left to right: without CI gas; with  $H_2$ ; with  $NH_3$ ; with  $H_2O$ ; and with HCl.

#### **RESULTS AND DISCUSSION**

Figures 1-4 summarize the spectra observed under CI conditions for mixtures of  $C_{60}$  with  $H_2$ ,  $N_3$ ,  $H_2O$  and HCl (Fig. 1), CH<sub>4</sub> (Fig. 2),  $C_4H_4$ (Fig. 3) and *i*- $C_4H_{10}$  (Fig. 4). The spectra begin at around m/z 700. The spectra between m/z 20 and m/z 700 are not given since they showed only the peaks expected from the CI reagent gas. One feature of the spectra shown in Figs. 1-4 which is immediately striking is the relatively high abundance of protonated  $C_{60}$ ,  $C_{60}H^{++}$ , in all spectra except that with ammonia. With HCl,  $C_{60}H^{+}$  clearly dominates the spectrum. A second, perhaps more noteworthy feature of these spectra is the abundance of higher mass-satellite peaks, particularly with  $H_2$ , NH<sub>3</sub>, CH<sub>4</sub>,  $C_2H_4$  and *i*- $C_4H_{10}$ . The spectrum in methane is particularly remarkable in this regard.

A more detailed inspection of the spectra shown in Figs. 1–4 reveals still further interesting features:

(1) The "self-CI" spectrum of  $C_{60}$  shows no surprises. The  $C_{60}^{+}$  peak predominates and, interestingly, there was no evidence for the formation of the dimer ion  $(C_{60})_2^{+}$ .



Fig. 2. CI spectrum of  $C_{60}$  with methane.

(2) Hydrogenation of  $C_{60}^{++}$  with up to at least 37 hydrogen atoms is observed in hydrogen with apparent special stabilities for  $C_{60}H_5^+$  and  $C_{60}H_9^+$  and perhaps also for  $C_{60}H_{23}^+$  and  $C_{60}H_{37}^+$ .

(3) The ammonia spectrum shows extensive hydrogenation. The strongest peaks correspond to  $C_{60}^{+}$  hydrogenated with an odd number of hydrogen atoms, viz.  $C_{60}H_x^+$  with x = 3, 5, 7, 9, 11 and 13. Values of x = 5, 7 and 9 particularly seem to be preferred. The peaks at  $[C_{60} + 19, 20, 21]^+$  very likely reflect additional substitution by nitrogen.

(4) Aside from the presence of  $C_6H^+$ , there is little further evidence for the hydrogenation of  $C_{60}^{,+}$  in the water spectrum. The small peaks at  $[C_{60} + 16]^+$  and  $[C_{60} + 17]^+$  may correspond to  $C_{60}O^{,+}$  and  $C_{60}OH^+$  respectively.

(5) Chlorine derivatives of  $C_{60}^{+}$  are noticeably absent in the HCl spectrum which is dominated by  $C_{60}H^+$ .

(6) The CI spectrum in methane/ $C_{60}$  shows a complicated pattern of peaks corresponding to  $C_{60}^{+}$  derivatized by  $nCH_x$  units. The spectrum appears to peak at about n = 14 and extends to about n = 32. Interestingly, it is quite similar to the fast atom bombardment (FAB) mass spectrum reported by Krusic et al. [16b] for the products obtained by UV irradiation of a benzene



Fig. 3. CI spectrum of  $C_{60}$  with ethylene.

solution of C<sub>60</sub> and di-*tert*-butylperoxide at 100°C that shows the formation of C<sub>60</sub>(CH<sub>3</sub>)<sup>+</sup><sub>n</sub> adducts with n = 1-34.

(7) The ethylene/ $C_{60}$  spectrum also shows derivative peaks corresponding to  $nCH_x$  to  $C_{60}^{+}$  but these peaks are much less intense relative to the  $C_{60}^{+}$  and  $C_{60}H^+$  peak heights and are readily discernible only up to approximately n = 6. The peaks at  $[M + 32, 60 \text{ and } 88]^+$  appear to reflect a series of  $C_2H_4$  additions.

(8) The spectrum with *iso*-butane also shows the pattern corresponding to the addition of  $CH_x$  units, in this case extending to at least n = 11.

# Formation of $C_{60}H^+$

Nothing is known about the reactivity of the radical cations of the molecules used as CI reagents in this study toward  $C_{60}$ . Table 1 shows that their relatively high recombination energies and low deprotonation energies make both charge transfer and proton transfer thermodynamically possible as shown in reaction 1. The ionization energy of  $C_{60}$  is only 7.61 eV [10], but the





Fig. 4. CI spectrum of C<sub>60</sub> with iso-butane.

### TABLE 1

Thermochemical properties of gases used as CI reagents<sup>a</sup>

Molecule	IE (eV)	PA(XH)	PA(X)	HA(X)	HA(X <sup>+</sup> )
H,	15.43	101.3	62	104	62
CH₄	12.51	132.0	130	105	43
нсі	12.75	134.8	128	103	108
C <sub>2</sub> H₄	10.51	162.6	174	103	63
$i-C_4H_{10}$	10.57	163.3	170	100	39
H <sub>2</sub> O	12.61	166.5	142	119	128
NH <sub>3</sub>	10.16	204.0	178	108	131

<sup>a</sup> IE = ionization energy; PA = proton affinity; HA = hydrogen atom affinity. Unless indicated otherwise, all values are in kcalmol<sup>-1</sup>. Proton affinities for XH were taken from the compilation of S.G. Lias, J.F. Liebman and R.D. Levin, J. Phys. Chem. Ref. Data, 13 (1984) 695. Other values were taken or derived from the compilation of S.G. Lias, J.E. Bartmess, J.F. Liebman, J.L. Holmes, R.D. Levin and W.G. Mallard, J. Phys. Chem. Ref. Data, 17 (1988) 1. proton affinity of  $C_{60}$  is very high,  $PA(C_{60}) = 204 \text{ kcal mol}^{-1}$  [19]. XH<sup>++</sup> +  $C_{40} \rightarrow C_{40}^{++} + \text{XH}$ 

$$XH^{+} + C_{60} \rightarrow C_{60}^{+} + XH$$
 (1b)  
 $\rightarrow C_{60}H^{+} + X^{-}$  (1b)

Of course the establishment of  $XH_2^+$  by self-protonation makes reactions of type 2 very likely sources of  $C_{60}H^+$ . Proton transfer reactions of this type have been observed with  $CH_5^+$ ,  $H_3O^+$  and  $NH_4^+$  [19]. The proton transfer reaction between  $NH_4^+$  and  $C_{60}$  is slow because of the similarity in the proton affinities of  $C_{60}$  and  $NH_3$  [19]; this may account for the relatively small  $[C_{60} + 1]^+$  signal in the ammonia spectrum.

$$XH_2^+ + C_{60} \rightarrow C_{60}H^+ + XH$$
 (2)

We can exclude hydrogen atom transfer reactions of type 3 with the CI reagent XH as a possible source of  $C_{60}$  H<sup>+</sup>. The hydrogen atom affinity of  $C_{60}^{*+}$  is very low, HA( $C_{60}^{*+}$ ) = 63 kcal mol<sup>-1</sup> [22], compared with HA(X) which has values exceeding 100 kcal mol<sup>-1</sup> for the CI reagents studied here (see Table 1). Indeed,  $C_{60}^{*+}$  has now been shown to be unreactive at 294 ± 2K toward H<sub>2</sub>, H<sub>2</sub>O, CH<sub>4</sub>, C<sub>2</sub>H<sub>4</sub> and *i*-C<sub>4</sub>H<sub>10</sub> and to react only slowly with NH<sub>3</sub> to produce the adduct ion  $C_{60}$ NH<sub>3</sub><sup>++</sup> [8,9].

$$C_{60}^{*+} + XH \to C_{60}H^+ + X^{*}$$
 (3)

Further hydrogenation of  $C_{60}^{,+}$ 

The extensive hydrogenation, with a preference for the formation of  $C_{60}^{++}$  with an odd number of hydrogen atoms, observed in the spectra with H<sub>2</sub> and NH<sub>3</sub> is most intriguing. Sequential hydrogenation reactions of type 4 are unlikely because of the relatively high hydrogen atom affinity of X.

$$C_{60}H_n^+ + XH \to C_{60}H_{n+1}^{+} + X$$
 (4)

Derivatization of  $C_{60}^{,+}$ 

The derivatization of  $C_{60}^{*+}$  observed with ammonia, methane, ethylene and *iso*-butane is also difficult to rationalize, particularly in view of the observed non-reactivity of  $C_{60}^{*+}$  toward methane, ethylene and isobutane at 294 ± 2 K [8] which rules out reactions of type 5. Addition reactions between  $C_{60}H_x^+$  (x = 2-4) and NH<sub>3</sub>, in analogy with the observed addition reaction between  $C_{60}^{*+}$  and ammonia [9], may account for the formation of the peaks at  $[C_{60} + 19, 20, 21]^+$ .

$$C_{60}^{'+} + XH \to C_{60}X^{+} + H^{'}$$
 (5)

The analogous reaction 6 in which the reacting XH is ionized instead of  $C_{60}$ 

is much more exothermic (by the difference in the ionization energies of Xand  $C_{60}$  which is at least 2.5 eV). Inspection of the hydrogen atom affinities of X<sup>+</sup> suggests that thermochemistry favours substitution with  $C_4H_9^+$ ,  $CH_3^+$ ,  $H^+$ and  $C_2H_3^+$  in reactions of type 6 and, more generally, type 7, both of which must compete with proton transfer

$$XH'' + C_{60} \rightarrow C_{60}X' + H'$$
 (6)

$$XH'' + C_{60}H_xX_{\nu} \to C_{60}H_xX_{\nu+1}^+ + H'$$
(7)

 $C_{60}H_xX_{y+1}^+$  ions are representative of the derivatized  $C_{60}^{,+}$  ions seen with  $CH_4$ ,  $C_2H_4$ , i- $C_4H_{10}$  and even  $H_2$ . Their formation depends on the availability of neutral molecules of type  $C_{60}H_xX_y$  whose origin is questionable. In the CI source they may be derived by electron/ion recombination of  $C_{60}X^+$ , where  $X = CH_3$ ,  $C_2H_3$  or  $C_4H_9$ , or perhaps by reactions of  $C_{60}X^+$  with radical species followed by electron/ion recombination. For example,  $C_{60}H^+$  is known to react unusually rapidly with H atoms under SIFT conditions to form  $C_{60}H_2^{++}$  [18].

# ACKNOWLEDGEMENTS

Financial support of our work by the Deutsche Forschungsgemeinschaft and the Fonds der Chemischen Industrie is greatly appreciated.

# NOTE ADDED IN PROOF

At the 40th ASMS Conference on Mass Spectrometry, 31 May–5 June 1992, Washington, DC, C.N. McEwan reported results using  $C_{60}$  as a radical sponge. He suggested that radical adducts generated under chemical ionization conditions are largely due to surface processes rather than ion/radical gas-phase reactions. Results of his experiments are also described in the 20 May 1992 issue of the Journal of the American Chemical Society.

#### REFERENCES

- 1 W. Krätschmer, L.D. Lamb, K. Fostiropoulos and D.R. Huffmann, Nature, 347 (1990) 354.
- 2 H.W. Kroto, J.R. Heath, S.C. O'Brien, R.F. Curl and R.E. Smalley, Nature, 318 (1985) 162.
- 3 (a) T.P. Snow and C.G. Seab, Astron. Astrophys., 213 (1989) 291.
  (b) W.B. Somerville and J.G. Bellis, Mon. Not. R. Astron. Soc., 240 (1989) 41P.
- 4 (a) A. Léger, L. d'Hendecourt, L. Verstraete and W. Schmidt, Astron. Astrophys., 203 (1988) 145.

(b) C. Joblin, J.P. Mailard, L. d'Hendecourt and A. Léger, Nature, 346 (1990) 729.

5 A. Webster, Nature, 352 (1991) 412.

- 6 H. Schwarz, Angew. Chem., Int. Ed. Engl., 31 (1992) 293.
- 7 (a) F.D. Weiss, J.L. Elkind, S.C. O'Brien, R.F. Curl and R.E. Smalley, J. Am. Chem. Soc., 110 (1988) 4464.
  (b) S.W. McElvany, H.H. Nelson, A.P. Baronavski, C.H. Watson and J.R. Eyler, Chem. Phys. Lett., 134 (1987) 214.
- 8 S. Petrie, G. Javahery, J. Wang and D.K. Bohme, J. Am. Chem. Soc., submitted.
- 9 G. Javahery, S. Petrie, J. Wang and D.K. Bohme, J. Am. Chem. Soc., submitted.
- 10 J.A. Zimmerman, J.R. Eyler, S.B.H. Bach and S.W. McElvany, J. Chem. Phys., 94 (1991) 3556.
- 11 R.E. Hauffler, J. Conceicao, P.F. Chibante, Y. Chai, N.E. Byrne, F. Flanagan, M.M. Haley, S.C. O'Brien, C. Pan, Z. Xiao, W.E. Billups, M.A. Ciufolini, R.H. Hauge, J.L. Margrave, L.J. Wilson, R.F. Curl and R.E. Smalley, J. Phys. Chem., 94 (1990) 8634.
- 12 H. Selig, C. Lifshitz, T. Peres, J.E. Fisher, A.R. McGie, W.J. Romanow, J.P. McCauley, Jr. and A.B. Smith III, J. Am. Chem. Soc., 113 (1991) 5475.
- 13 F. Diederich, R. Ettl, Y. Rubin, R.L. Whetten, R. Beck, M. Alvarez, S. Anz, D. Sehsharma, F. Wudl, K.C. Khemani and A. Koch, Science, 252 (1991) 548.
- 14 J.M. Wood, B. Kahr, S.H. Hoke II, L. Dejarme, R.G. Cooks and D. Ben-Amotz, J. Am. Chem. Soc., 113 (1991) 5907.
- 15 J.W. Bausch, G.K. Surja Prakash, G.A. Olah, D.S. Tse, D.C. Lorentz, Y.K. Bae and M. Ripudaman, J. Am. Chem. Soc., 113 (1991) 3205.
- (a) P.J. Krusic, E. Wasserman, B.A. Parkinson, B. Malone, E.R. Holler Jr., P.N. Keizer, J.R. Morton and K.F. Preston, J. Am. Chem. Soc., 113 (1991) 6274.
  (b) P.J. Krusic, E. Wasserman, P.N. Keizer, J.R. Morton and K.F. Preston, Science, 254 (1991) 1183.
- 17 T. Suzuki, Q. Li, K.C. Khemani, F. Wudl and Ö. Almarsson, Science, 254 (1991) 1186.
- 18 S. Petrie, G. Javahery, J. Wang and D.K. Bohme, J. Am. Chem. Soc., in press.
- 19 S.W. McElvany and J.H. Callahan, J. Phys. Chem., 95 (1991) 6187.
- 20 (a) L.S. Roth, Y. Huang, J.T. Schwedler, C.J. Cassidy, D. Ben-Amotz, B. Kahr and B.S. Freiser, J. Am. Chem. Soc., 113 (1991) 6198.
  - (b) Y. Huang and B.S. Freiser, J. Am. Chem. Soc., 113 (1991) 8186.
  - (c) Y. Huang and B.S. Freiser, J. Am. Chem. Soc., 113 (1991) 9418.
  - (d) G. Javahery, S. Petrie, J. Wang and D.K. Bohme, Chem. Phys. Lett., submitted.
- (a) R. Srinivas, D. Sülzle, T. Weiske and H. Schwarz, Int. J. Mass Spectrom. Ion Processes, 107 (1991) 368.
  (b) R. Srinivas, D. Sülzle, W. Koch, C.H. DePuy and H. Schwarz, J. Am. Chem. Soc., 113

(1991) 5970.

22 D.K. Bohme, Chem. Rev., submitted.